



Classification of High-resolution Solar Hα Spectra using t-SNE

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Why do we need classification?

□ On one observing day → time-series of 21 Hα spatio-spectral data cubes.



3D data -- Science in every pixel

□ Contains about 8.7 million intensity and contrast profiles.



t-SNE - Appropriate tool to classify spectra

- Probabilistic approach
- Dimensionality reduction
- t-SNE result of classifying on 3000 256-dimensional grayscale images of handwritten digits.
- Classes are quite well separated even though t-SNE had no information about class labels.
- Within each class, properties like orientation, skew and stroke thickness tend to vary smoothly across the space.

L.J.P. van der Maaten and G.E. Hinton. **Visualizing High-Dimensional Data Using t-SNE**. *Journal of Machine Learning Research* 9(Nov):2579-2605, 2008

L.J.P. van der Maaten. Accelerating t-SNE using Tree-Based Algorithms. *Journal of Machine Learning Research* 15(Oct):3221-3245, 2014.



t-SNE -- From profiles to classification



Q1 Is the default choice ok?

Q2 Is the projection different for profiles and PCA coefficients?

Q3 Is the projection affected by seeing conditions?

https://distill.pub/2016/misread-tsne/

t-SNE projection of 630 x 660 spectral profiles with 601 wavelength points.

The choice of parameters perplexity = 50, theta = 0.5, number of iterations = 1000

□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.3$, P = 50, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.4$, P = 50, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.7$, P = 50, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.8$, P = 50, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.5$, P = 10, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.5$, P = 30, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.5$, P = 50, n = 1000



□ Three parameters we can change.

Theta, Perplexity, Number of Iterations $\rightarrow \theta = 0.5$, P = 80, n = 1000















A1 The default parameters are fine, maybe the number of iteration has to be improved for large datasets



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Good or Bad Seeing PCA or Observed



A2 Seeing does affect the projection but not much A3 PCA coefficients and observed contrast profiles lead to similar results.

Back mapping



Selected the regions after using threshold of 0.9 and better.



These are the regions where the profiles can surely be inverted using cloud model.

Conclusion

- t-SNE is a powerful tool to classify spectra.
- □ No prior information is needed.
- It classify good vs. bad profiles for inversion.
- \Box Best settings perplexity = 50,
 - theta = 0.5 and number of iterations = 1000 → based on time for computation and discerning power.
- Contrast as well as line profiles,
 PCA coefficients, PCA denoised or observed profiles lead to similar projection.

- Does show some differences for good and bad seeing.
- The regions which can surely be inverted using cloud models are discernible.
- Performing cloud model inversions of selected regions in t-SNE projection.
- Projecting more data points on the already projected map.
- □ Comparison with UMAP.